

Majorana CP-Violation, $(\beta\beta)_{0\nu}$ -Decay and Leptogenesis

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NOW'06
Conca Specchiulla, Otranto
September 9 - 16, 2006

Compelling Evidences for ν -Oscillations

– ν_{atm} : **SK** UP-DOWN ASYMMETRY

θ_{23} -, L/E - dependences of μ -like events

Dominant $\nu_{\mu} \rightarrow \nu_{\tau}$ K2K; MINOS, CNGS (OPERA)

– ν_{\odot} : **Homestake, Kamiokande, SAGE, GALLEX/GNO**

Super-Kamiokande, SNO; KamLAND

Dominant $\nu_e \rightarrow \nu_{\mu, \tau}$ **BOREXINO, ..., LowNu**

– **LSND**

Dominant $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ **MiniBOONE**

$$\nu_{lL} = \sum_{j=1} U_{lj} \nu_{jL} \quad l = e, \mu, \tau. \quad (1)$$

Three Neutrino Mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} . \quad (2)$$

U is the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) neutrino mixing matrix,

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \quad (3)$$

- U - $n \times n$ unitary:

	n	2	3	4
mixing angles:	$\frac{1}{2}n(n-1)$	1	3	6

CP-violating phases:

ν_j - Dirac:	$\frac{1}{2}(n-1)(n-2)$	0	1	3
ν_j - Majorana:	$\frac{1}{2}n(n-1)$	1	3	6

$n = 3$: 1 Dirac and
2 additional CP-violating phases, Majorana phases

PMNS Matrix: Standard Parametrization

$$U = V \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix} \quad (4)$$

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13}e^{i\delta} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13}e^{i\delta} \end{pmatrix} \quad (5)$$

- $s_{ij} \equiv \sin \theta_{ij}$, $c_{ij} \equiv \cos \theta_{ij}$, $\theta_{ij} = [0, \frac{\pi}{2}]$,
- δ - Dirac CP-violation phase, $\delta = [0, 2\pi]$,
- α_{21} , α_{31} - the two Majorana CP-violation phases.
- $\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 \cong 8.0 \times 10^{-5} \text{ eV}^2 > 0$, $\sin^2 \theta_{12} \cong 0.30$, $\cos 2\theta_{12} \gtrsim 0.28$ (2σ),
- $|\Delta m_{\text{atm}}^2| \equiv |\Delta m_{31}^2| \cong 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{23} \cong 1$,
- θ_{13} - the CHOOZ angle: $\sin^2 \theta_{13} < 0.027$ (0.041) 2σ (3σ).

A.Bandyopadhyay, S.Choubey, S.Goswami, S.T.P., D.P.Roy, hep-ph/0406328 (updated)

T. Schwetz, hep-ph/0606060.

- $\sqrt{\Delta m_{\odot}^2} \sin^2 \theta_{12} \cong 3.0 \times 10^{-3} \text{ eV}$ (\pm) $\sqrt{|\Delta m_{\text{atm}}^2|} \sin^2 \theta_{13} \lesssim 2.2 \times 10^{-3} \text{ eV}$;
- $\sqrt{|\Delta m_{\text{atm}}^2|} \cong 5 \times 10^{-2} \text{ eV}$; $\sqrt{|\Delta m_{\text{atm}}^2|} \cos 2\theta_{12} \gtrsim 1.4 \times 10^{-2} \text{ eV}$ ($\cos 2\theta_{12} \gtrsim 0.28$)
- m_0 : $m_0^2 \gg \Delta m_{\odot}^2, |\Delta m_{\text{atm}}^2|$, $m_0 \gtrsim 0.1 \text{ eV}$
- $\text{sgn}(\Delta m_{\text{atm}}^2) = \text{sgn}(\Delta m_{31}^2)$ not determined

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \text{ normal mass ordering}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \text{ inverted mass ordering}$$

Convention: $m_1 < m_2 < m_3$ - **NMO**, $m_3 < m_1 < m_2$ - **IMO**

- **Majorana phases α_{21}, α_{31} :**

- $\nu_l \leftrightarrow \nu_l, \bar{\nu}_l \leftrightarrow \bar{\nu}_l$ not sensitive;

S.M. Bilenky, J. Hosek, S.T.P., 1980;
P. Langacker, S.T.P., G. Steigman, S. Toshev, 1987

- $|\langle m \rangle|$ in $(\beta\beta)_{0\nu}$ -decay depends on α_{21}, α_{31} ;
- $\Gamma(\mu \rightarrow e + \gamma)$ etc. in SUSY theories depend on $\alpha_{21,31}$;
- BAU, leptogenesis scenario: $\alpha_{21,31}$!

Future Progress

- Determination of the nature - Dirac or Majorana, of ν_j .
- Determination of $\text{sgn}(\Delta m_{\text{atm}}^2)$, type of ν - mass spectrum

$$m_1 \ll m_2 \ll m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.10 \text{ eV.}$$

- Determining, or obtaining significant constraints on, the absolute scale of ν_j -masses, or $\min(m_j)$.
- Status of the CP-symmetry in the lepton sector: violated due to δ (Dirac), and/or due to α_{21}, α_{31} (Majorana)?
- High precision determination of $\Delta m_{\odot}^2, \theta_{\odot}, \Delta m_{\text{atm}}^2, \theta_{\text{atm}}$.
- Measurement of, or improving by at least a factor of (5 - 10) the existing upper limit on, $\sin^2 \theta_{13}$.
- Searching for possible manifestations, other than ν_l -oscillations, of the non-conservation of $L_l, l = e, \mu, \tau$, such as $\mu \rightarrow e + \gamma, \tau \rightarrow \mu + \gamma$, etc. decays.

- Understanding at fundamental level the mechanism giving rise to the ν - masses and mixing and to the L_l -non-conservation. Includes understanding
 - the origin of the observed patterns of ν -mixing and ν -masses ;
 - the physical origin of CPV phases in U_{PMNS} ;
 - Are the observed patterns of ν -mixing and of $\Delta m_{21,31}^2$ related to the existence of a new symmetry?
 - Is there any relations between q -mixing and ν - mixing? Is $\theta_{12} + \theta_c = \pi/4$?
 - Is $\theta_{23} = \pi/4$, or $\theta_{23} > \pi/4$ or else $\theta_{23} < \pi/4$?
 - Is there any correlation between the values of CPV phases and of mixing angles in U_{PMNS} ?
- Progress in the theory of ν -mixing might lead to a better understanding of the origin of the BAU.
 - Can the Majorana and/or Dirac CPVP in U_{PMNS} be the leptogenesis CPV parameters at the origin of BAU?

$(\beta\beta)_{0\nu}$ –Decay Experiments:

- Majorana nature of ν_j
- Type of ν –mass spectrum (NH, IH, QD)
- Absolute neutrino mass scale

${}^3\text{H}$ β -decay , cosmology: m_{ν} (QD, IH)

- CPV due to Majorana CPV phases

ν_j – Dirac or Majorana particles, fundamental problem

ν_j –Dirac: **conserved lepton charge exists**, $L = L_e + L_\mu + L_\tau$, $\nu_j \neq \bar{\nu}_j$

ν_j –Majorana: **no lepton charge is exactly conserved**, $\nu_j \equiv \bar{\nu}_j$

The observed patterns of ν –mixing and of Δm_{atm}^2 and Δm_{\odot}^2 can be related to Majorana ν_j and an **approximate** symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

See-saw mechanism: ν_j – Majorana

Establishing that ν_j are Majorana particles would be as important as the discovery of ν – oscillations.

If ν_j – Majorana particles, U_{PMNS} contains (3- ν mixing)

δ -Dirac, α_{21} , α_{31} - Majorana physical CPV phases

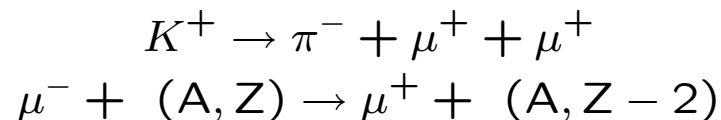
ν -oscillations $\nu_l \leftrightarrow \nu_{l'}$, $\bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$, $l, l' = e, \mu, \tau$,

- are not sensitive to the nature of ν_j ,

S.M. Bilenky et al., 1980;
P. Langacker et al., 1987

- provide information on $\Delta m_{jk}^2 = m_j^2 - m_k^2$, but not on the absolute values of ν_j masses.

The Majorana nature of ν_j can manifest itself in the existence of $\Delta L = \pm 2$ processes:



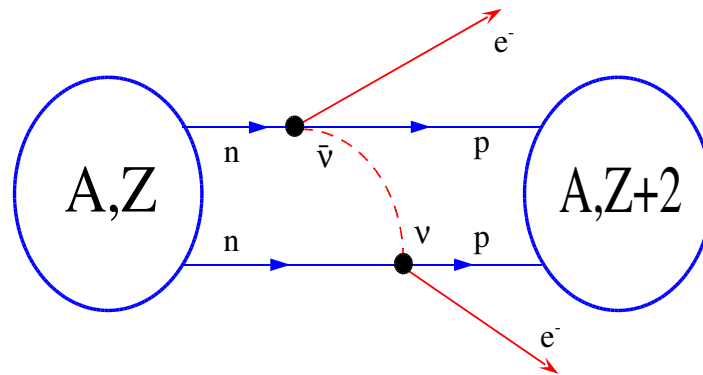
The process most sensitive to the possible Majorana nature of ν_j - $(\beta\beta)_{0\nu}$ -decay



of even-even nuclei, ^{48}Ca , ^{76}Ge , ^{82}Se , ^{100}Mo , ^{116}Cd , ^{130}Te , ^{136}Xe , ^{150}Nd .

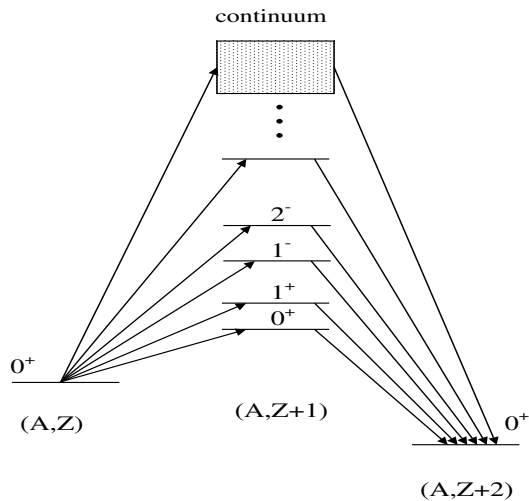
$2n$ from (A, Z) exchange a virtual Majorana ν_j (via the CC weak interaction) and transform into $2p$ of $(A, Z+2)$ and two free e^- .

Nuclear $0\nu\beta\beta$ -decay



strong in-medium modification of the basic process

$$dd \rightarrow uue^-e^-(\bar{\nu}_e\bar{\nu}_e)$$



virtual excitation
of states of all multiplicities
in $(A, Z+1)$ nucleus

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$\begin{aligned} |\langle m \rangle| &= |m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha_{21}} + m_3 |U_{e3}|^2 e^{i\alpha_{31}}| \\ &= |m_1 c_{12}^2 c_{13}^2 + m_2 s_{12}^2 c_{13}^2 e^{i\alpha_{21}} + m_3 s_{13}^2 e^{i\alpha_{31}}|, \quad \theta_{12} \equiv \theta_{\odot}, \theta_{13} - \text{CHOOZ} \end{aligned}$$

α_{21}, α_{31} - the two Majorana CPVP of the PMNS matrix.

CP-invariance: $\alpha_{21} = 0, \pm\pi, \alpha_{31} = 0, \pm\pi;$

$$\eta_{21} \equiv e^{i\alpha_{21}} = \pm 1, \quad \eta_{31} \equiv e^{i\alpha_{31}} = \pm 1$$

relative CP-parities of ν_1 and ν_2 , and of ν_1 and ν_3 .

L. Wolfenstein, 1981;

S.M. Bilenky, N. Nedelcheva, S.T.P., 1984;

B. Kayser, 1984.

Best sensitivity: Heidelberg-Moscow ^{76}Ge experiment.

Claim for a positive signal at $> 3\sigma$:

H. Klapdor-Kleingrothaus et al., PL B586 (2004),

$|\langle m \rangle| = (0.1 - 0.9) \text{ eV}$ (99.73% C.L.).

IGEX ^{76}Ge : $|\langle m \rangle| < (0.33 - 1.35) \text{ eV}$ (90% C.L.).

Taking data - NEMO3 (^{100}Mo), CUORICINO (^{130}Te):

$|\langle m \rangle| < (0.7 - 1.2) \text{ eV}$, $|\langle m \rangle| < (0.18 - 0.90) \text{ eV}$ (90% C.L.).

Large number of projects: $|\langle m \rangle| \sim (0.01 - 0.05) \text{ eV}$

CUORE - ^{130}Te ,

GERDA - ^{76}Ge ,

SuperNEMO - ^{82}Se ,

EXO - ^{136}Xe ,

MAJORANA - ^{76}Ge ,

MOON - ^{100}Mo ,

CANDLES - ^{48}Ca ,

XMASS - ^{136}Xe .

$$|\langle m \rangle| : m_j, \theta_\odot \equiv \theta_{12}, \theta_{13}, \alpha_{21,31}$$

$m_{1,2,3}$ - in terms of $\min(m_j)$, Δm_{atm}^2 , Δm_\odot^2

S.T.P., A.Yu. Smirnov, 1994

Convention: $m_1 < m_2 < m_3$ - **NMO**, $m_3 < m_1 < m_2$ - **IMO**

$$\Delta m_\odot^2 \equiv \Delta m_{21}^2, \quad m_2 = \sqrt{m_1^2 + \Delta m_\odot^2},$$

while either

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \quad m_3 = \sqrt{m_1^2 + \Delta m_{\text{atm}}^2}, \quad \text{normal mass ordering, or}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \quad m_1 = \sqrt{m_3^2 + |\Delta m_{\text{atm}}^2| - \Delta m_\odot^2}, \quad \text{inverted mass ordering}$$

The neutrino mass spectrum –

Normal hierarchical (NH) if $m_1 \ll m_2 \ll m_3$,

Inverted hierarchical (IH) if $m_3 \ll m_1 \cong m_2$,

Quasi-degenerate (QD) if $m_1 \cong m_2 \cong m_3 = m$, $m_j^2 \gg |\Delta m_{\text{atm}}^2|$; $m_j \gtrsim 0.1$ eV

Given $|\Delta m_{\text{atm}}^2|$, Δm_\odot^2 , θ_\odot , θ_{13} ,

$$|\langle m \rangle| = |\langle m \rangle| (m_{\text{min}}, \alpha_{21}, \alpha_{31}; S), \quad S = \text{NO(NH), IO(IH)}.$$

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$|\langle m \rangle| \cong \left| \sqrt{\Delta m_{\odot}^2} \sin^2 \theta_{12} e^{i\alpha} + \sqrt{\Delta m_{31}^2} \sin^2 \theta_{13} e^{i\beta_M} \right|, \quad m_1 \ll m_2 \ll m_3 \text{ (NH)},$$

$$|\langle m \rangle| \cong \sqrt{m_3^2 + \Delta m_{13}^2} |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_3 < (\ll) m_1 < m_2 \text{ (IH)},$$

$$|\langle m \rangle| \cong m |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_{1,2,3} \cong m \gtrsim 0.10 \text{ eV (QD)},$$

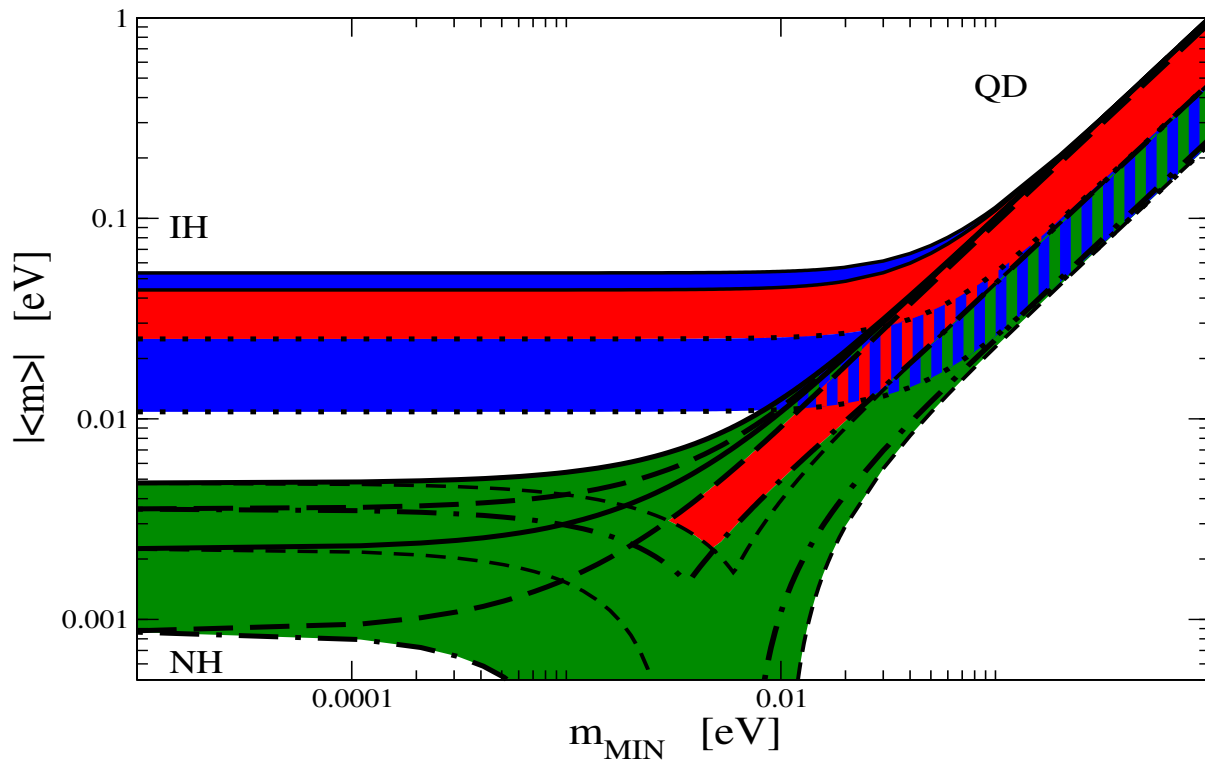
$$\theta_{12} \equiv \theta_{\odot}, \theta_{13}\text{-CHOOZ}; \quad \alpha \equiv \alpha_{21}, \beta_M \equiv \alpha_{31}.$$

$$\text{CP-invariance: } \alpha = 0, \pm\pi, \beta_M = 0, \pm\pi;$$

$$|\langle m \rangle| \lesssim 5 \times 10^{-3} \text{ eV, NH};$$

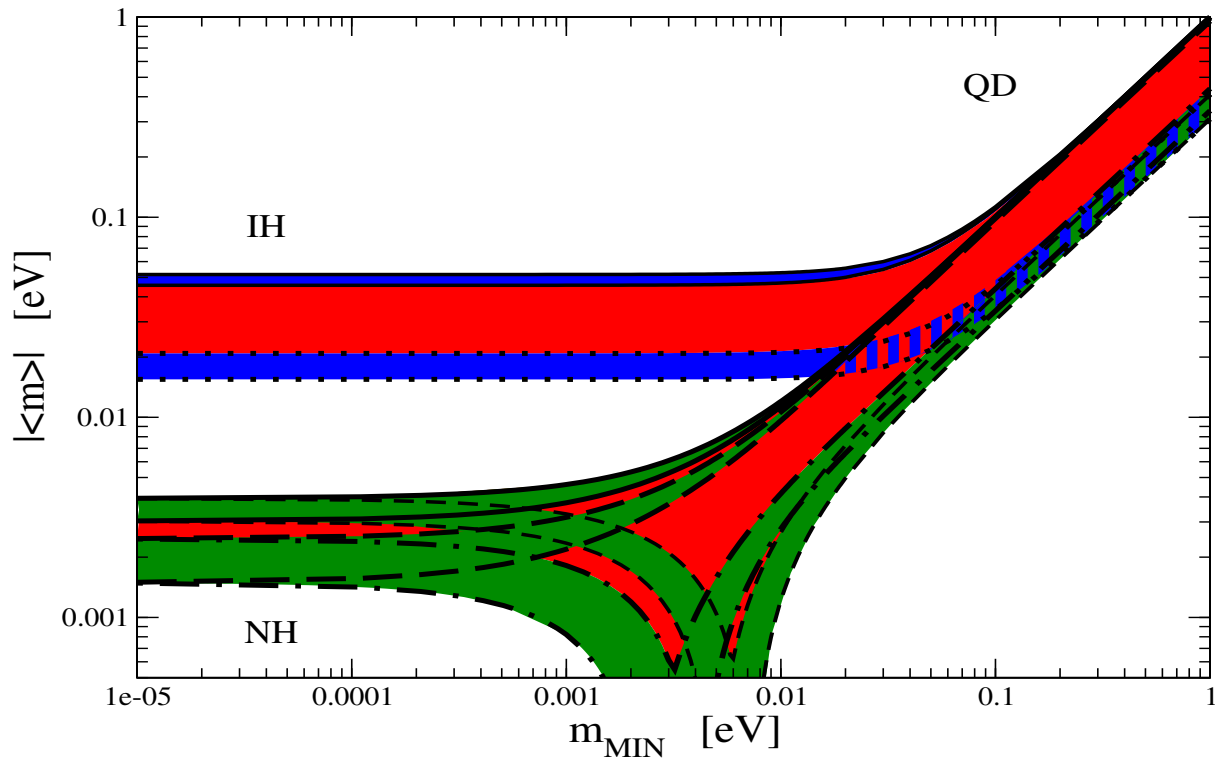
$$\sqrt{\Delta m_{13}^2} \cos 2\theta_{12} \cong 0.013 \text{ eV} \lesssim |\langle m \rangle| \lesssim \sqrt{\Delta m_{13}^2} \cong 0.055 \text{ eV, IH};$$

$$m \cos 2\theta_{12} \lesssim |\langle m \rangle| \lesssim m, \quad m \gtrsim 0.10 \text{ eV, QD}.$$



S. Pascoli, S.T.P., 2006

The current 2σ ranges of values of the parameters used.



$\sin^2 \theta_{13} = 0.015 \pm 0.006$; $1\sigma(\Delta m_{\odot}^2) = 4\%$, $1\sigma(\sin^2 \theta_{\odot}) = 4\%$, $1\sigma(|\Delta m_{\text{atm}}^2|) = 6\%$;

$2\sigma(|\langle m \rangle|)$ used.

Nuclear Matrix Element Uncertainty

$$|\langle m \rangle| = \zeta ((|\langle m \rangle|_{\text{exp}})_{\text{MIN}} \pm \Delta) , \quad \zeta \geq 1,$$

$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}}$ – obtained with the **maximal physically allowed value of NME**.

A measurement of the $(\beta\beta)_{0\nu}$ -decay half-life time

$$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}} - \Delta \leq |\langle m \rangle| \leq \zeta ((|\langle m \rangle|_{\text{exp}})_{\text{MIN}} + \Delta) .$$

The estimated range of ζ^2 :

$${}^{48}\text{Ca}, \quad \zeta^2 \simeq 3.5$$

$${}^{76}\text{Ge}, \quad \zeta^2 \simeq 10$$

$${}^{82}\text{Se}, \quad \zeta^2 \simeq 10$$

$${}^{130}\text{Te}, \quad \zeta^2 \simeq 38.7$$

S. Elliot, P. Vogel, 2002

NH vs IH (QD):

$$\zeta |\langle m \rangle|_{\text{max}}^{\text{NH}} < |\langle m \rangle|_{\text{min}}^{\text{IH(QD)}} , \quad \zeta \geq 1 .$$

IH vs QD:

$$\zeta |\langle m \rangle|_{\text{max}}^{\text{IH}} < |\langle m \rangle|_{\text{min}}^{\text{QD}} , \quad \zeta \geq 1 .$$

Method of Analysis

$$\Gamma_{\text{th}} = G |\mathcal{M}|^2 (|\langle m \rangle|(\mathbf{x}))^2, \quad \mathbf{x} = (\mathbf{x}_{\text{osc}}, \mathbf{x}_{\beta\beta}^{0\nu})$$

$$\mathbf{x}_{\text{osc}} = (\theta_{12}, \theta_{13}, |\Delta\mathbf{m}_{31}^2|, \Delta\mathbf{m}_{21}^2),$$

$$\mathbf{x}_{\beta\beta}^{0\nu} = (m_0, \text{sgn}(\Delta\mathbf{m}_{31}^2), \alpha_{21}, \alpha_{31}).$$

$$|\langle m \rangle|^{\text{obs}} \equiv \sqrt{\frac{\Gamma_{\text{obs}}}{G}} \frac{1}{|\mathcal{M}_0|}, \quad \sigma_{\beta\beta} = \frac{1}{2} \frac{1}{\sqrt{\Gamma_{\text{obs}} G}} \frac{1}{|\mathcal{M}_0|} \sigma(\Gamma_{\text{obs}}),$$

$|\mathcal{M}_0|$ is some nominal value of the NME.

$$\chi^2(\mathbf{x}_{\beta\beta}^{0\nu}, \mathbf{F}) = \min_{\xi \in [1/\sqrt{F}, \sqrt{F}]} \frac{[\xi |\langle m \rangle|(\mathbf{x}) - |\langle m \rangle|^{\text{obs}}]^2}{\sigma_{\beta\beta}^2 + \xi^2 \sigma_{\text{th}}^2}.$$

$$\xi \equiv \frac{|\mathcal{M}|}{|\mathcal{M}_0|}, \quad \xi = [1/\sqrt{F}, \sqrt{F}], \quad F \geq 1,$$

$|\mathcal{M}|$ is the *true* value of the NME.

Majorana CPV Phases and $|\langle m \rangle|$

IH spectrum: $m_{\min} < 0.01$ eV, $\sin^2 \theta$ – negligible

$$\sqrt{\Delta m_{\text{atm}}^2} |\cos 2\theta_{\odot}| \leq |\langle m \rangle| \leq \sqrt{\Delta m_{\text{atm}}^2}.$$

“Just CP-violating” region:

$$(|\langle m \rangle|_{\text{exp}})_{\text{MAX}} < \sqrt{(\Delta m_{\text{atm}}^2)_{\text{MIN}}},$$

$$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}} > \sqrt{(\Delta m_{\text{atm}}^2)_{\text{MAX}} (\cos 2\theta_{\odot})_{\text{MAX}}},$$

$$|\langle m \rangle| = \zeta (|\langle m \rangle|_{\text{exp}})_{\text{MIN}} \pm \Delta, \quad \zeta \geq 1$$

Necessary condition for establishing CP-violation:

$$1 \leq \zeta < \frac{\sqrt{(\Delta m_{\text{atm}}^2)_{\text{MIN}}}}{\sqrt{(\Delta m_{\text{atm}}^2)_{\text{MAX}} (\cos 2\theta_{\odot})_{\text{MAX}} + 2\Delta}} \simeq \frac{1}{(\cos 2\theta_{\odot})_{\text{MAX}}}$$

QD spectrum, $m_{1,2,3} \simeq m_0 \gtrsim 0.20$ eV - similar condition: $\Delta m_{\text{atm}}^2 \rightarrow m_0^2$.

CPV can be established provided

- $|\langle m \rangle|$ measured with $\Delta \lesssim 15\%$;
- Δm_{atm}^2 (IH) or m_0 (QD) measured with $\delta \lesssim 10\%$;
- $\xi \lesssim 1.5$;
- α_{21} (QD): in the interval $\sim [\frac{\pi}{4} - \frac{3\pi}{4}]$, or $\sim [\frac{5\pi}{4} - \frac{3\pi}{2}]$;
- $\tan^2 \theta_{\odot} \gtrsim 0.40$.

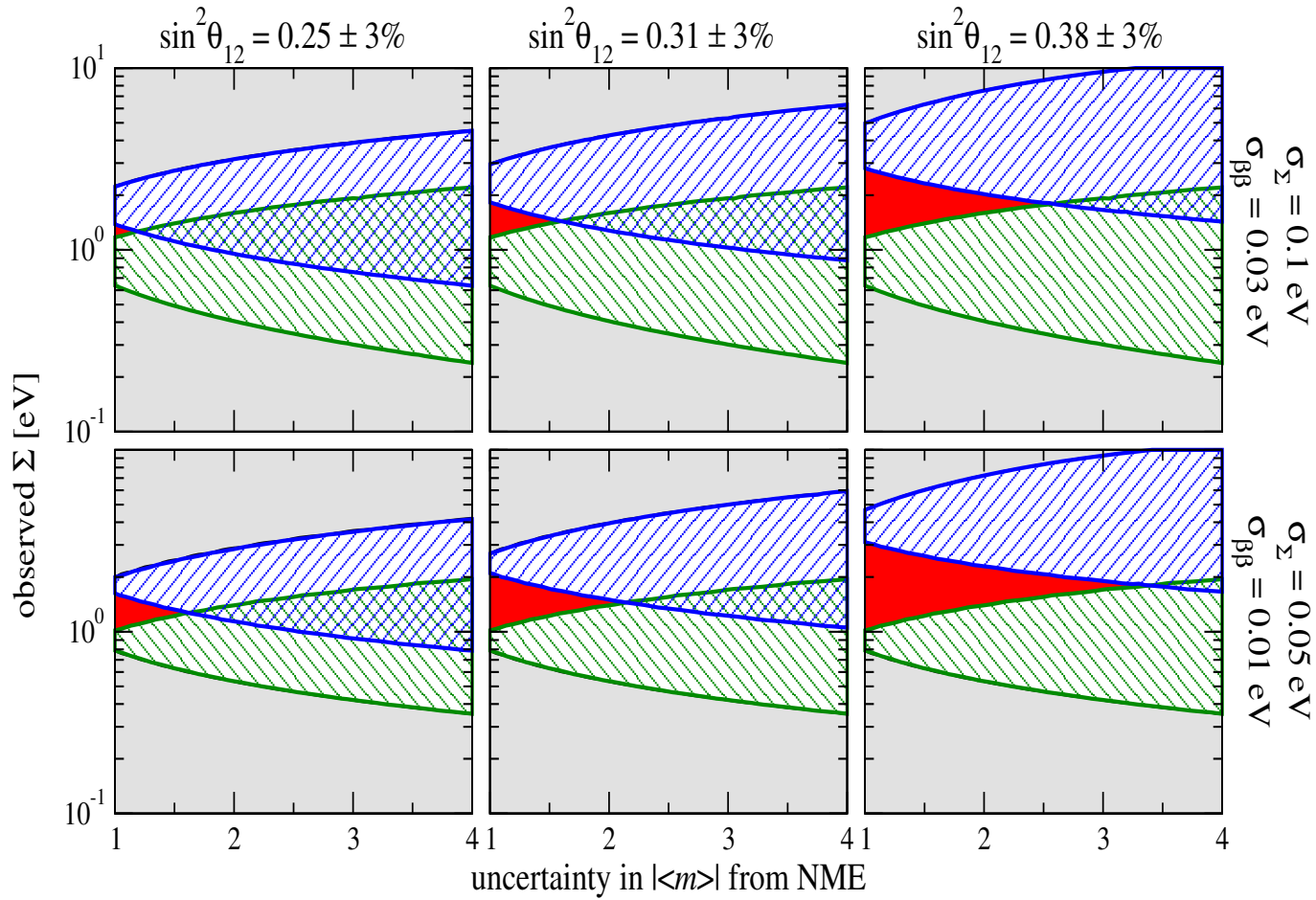
S. Pascoli, S.T.P., W. Rodejohann, 2002

S. Pascoli, S.T.P., L. Wolfenstein, 2002

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226


No “No-go for detecting CP-Violation via $(\beta\beta)_{0\nu}$ -decay”

V. Barger *et al.*, 2002



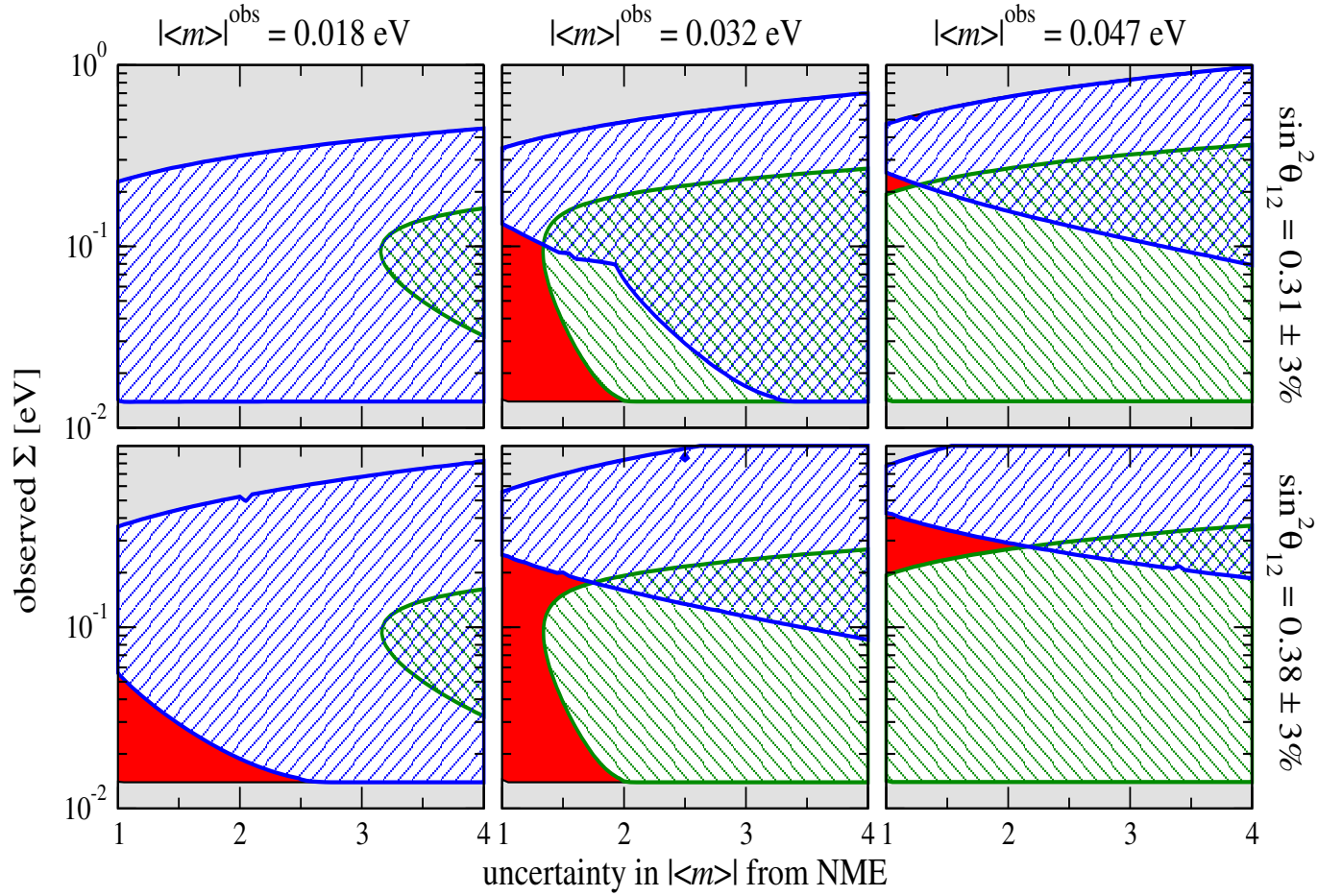
 data consistent with $\alpha_{21} = \pi$

 data consistent with $\alpha_{21} = 0$


 $\langle m \rangle$ and Σ inconsistent at 2σ


 CP violation established at 2σ

$\sin^2 \theta_{13} = 0 \pm 0.002$, $\Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%$, $\Delta m_{31}^2 = 2.2 \times 10^{-3} \pm 3\%$ observed $\langle m \rangle = 0.3 \text{ eV}$



 data consistent with $\alpha_{21} = \pi$

 data consistent with $\alpha_{21} = 0$

 $\langle m \rangle$ and Σ inconsistent at 2σ

 CP violation established at 2σ

$$\sin^2 \theta_{13} = 0 \pm 0.002, \quad \Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%, \quad \Delta m_{31}^2 = -2.2 \times 10^{-3} \pm 3\%, \quad \sigma_{\beta\beta} = 0.004 \text{ eV}, \quad \sigma_{\Sigma} = 0.04 \text{ eV}$$

Absolute Neutrino Mass Measurements

The Troitzk and Mainz ${}^3\text{H}$ β -decay experiments

$$m_{\nu_e} < 2.3 \text{ eV} \quad (95\% \text{ C.L.})$$

There are prospects to reach sensitivity

$$\text{KATRIN :} \quad m_{\nu_e} \sim 0.2 \text{ eV}$$

Cosmological and astrophysical data: the WMAP result combined with data from large scale structure surveys (2dFGRS, SDSS)

$$\sum_j m_j \equiv \Sigma < (0.4 - 1.7) \text{ eV}$$

The WMAP and future PLANCK experiments can be sensitive to

$$\sum_j m_j \cong 0.4 \text{ eV}$$

Data on weak lensing of galaxies by large scale structure, combined with data from the WMAP and PLANCK experiments may allow to determine

$$\sum_j m_j : \quad \delta \cong 0.04 \text{ eV.}$$

M_ν from the See-Saw Mechanism

P. Minkowski, 1977.

M. Gell-Mann, P. Ramond, R. Slansky, 1979;

T. Yanagida, 1979;

R. Mohapatra, G. Senjanovic, 1980.

- Explains the smallness of ν -masses.
- Through **leptogenesis theory** links the ν -mass generation to the generation of baryon asymmetry of the Universe Y_B .

S. Fukugita, T. Yanagida, 1986.

- In SUSY GUT's with see-saw mechanism of ν -mass generation, the LFV decays

$$\mu \rightarrow e + \gamma, \quad \tau \rightarrow \mu + \gamma, \quad \tau \rightarrow e + \gamma, \quad \text{etc.}$$

are predicted to take place with rates within the reach of present and future experiments.

F. Borzumati, A. Masiero, 1986.

- The ν_j are **Majorana particles**; $(\beta\beta)_{0\nu}$ -decay is allowed.

See-Saw: Dirac ν -mass m_D + Majorana mass M_R for N_R

The Role of the CPV Phases

$$-\mathcal{L} = \overline{N_{Ri}} (m_D)_{ij} \nu_{Lj} + \frac{1}{2} \overline{(N_{Ri})^c} (M_R)_{ij} N_{Rj}$$

m_D generated by the Yukawa interaction:

$$-\mathcal{L}_Y = \overline{N_{Ri}} (Y_\nu)_{ij} L_j H_u, \quad v_u = 174 \text{ GeV} \sin \beta, \quad v_u Y_\nu = m_D - \text{complex}$$

For M_R - sufficiently large,

$$m_\nu \simeq m_D^T M_R^{-1} m_D = v_u^2 Y_\nu^T M_R^{-1} Y_\nu = U_{\text{PMNS}}^* m_\nu^{\text{diag}} U_{\text{PMNS}}^\dagger.$$

Basis: $M_R = (M_1, M_2, M_3)$; $D_N \equiv \text{diag}(M_1, M_2, M_3)$, $D_\nu \equiv \text{diag}(m_1, m_2, m_3)$.

$Y_\nu = \sqrt{D_N} R \sqrt{D_\nu} (U_{\text{PMNS}})^\dagger / v_u$, all at M_R ; R -complex, $R^T R = \mathbf{1}$.

J.A. Casas and A. Ibarra, 2001

In GUTs, $M_R < M_X$, $M_X \sim 10^{16}$ GeV;

in GUTs, e.g., $M_R = (10^9, 10^{12}, 10^{15})$ GeV, $m_D \sim 1$ GeV.

Leptogenesis

$$Y_B = \frac{n_B - n_{\bar{B}}}{S} \sim 6 \times 10^{-10}$$

$$Y_B \cong -10^{-2} \quad \kappa \varepsilon$$

W. Buchmüller, M. Plümacher, 1998;

W. Buchmüller, P. Di Bari, M. Plümacher, 2004

κ - efficiency factor; $\kappa \sim 10^{-1} - 10^{-3}$: $\varepsilon \gtrsim 10^{-7}$.

ε : CP -, L - violating asymmetry generated in out of equilibrium N_{Rj} -decays in the early Universe,

$$\varepsilon_1 = \frac{\Gamma(N_1 \rightarrow \Phi^- \ell^+) - \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}{\Gamma(N_1 \rightarrow \Phi^- \ell^+) + \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}$$

$$\simeq \frac{1}{8\pi} \frac{1}{(Y_\nu Y_\nu^\dagger)_{11}} \sum_{j=2,3} \text{Im}(Y_\nu Y_\nu^\dagger)_{1j}^2 (f(M_j^2/M_1^2) + g(M_j^2/M_1^2)) .$$

$$f(x) = \sqrt{x} \left(1 - (1+x) \ln \left(\frac{1+x}{x} \right) \right) , \quad g(x) = \frac{\sqrt{x}}{1-x}$$

M.A. Luty, 1992;

L. Covi, E. Roulet, and F. Vissani, 1996

$$\frac{1}{\kappa} \simeq \frac{3.3 \times 10^{-3} \text{ eV}}{\tilde{m}_1} + \left(\frac{\tilde{m}_1}{0.55 \times 10^{-3} \text{ eV}} \right)^{1.16} , \quad \tilde{m}_1 \equiv \frac{v_u^2}{M_1} (Y_\nu Y_\nu^\dagger)_{11}$$

G. F. Giudice et al., 2004

LFV Charged Lepton Decays

$m_\nu \neq 0$, $U_{\text{PMNS}} \neq \mathbf{1}$: L_e , L_μ , L_τ not conserved

$\mu \rightarrow e + \gamma$, $\mu \rightarrow 3e$, $\tau \rightarrow \mu + \gamma$, $\tau \rightarrow e + \gamma$, etc., allowed.

$$BR(\mu \rightarrow e + \gamma) < 1.2 \times 10^{-11} \quad (\text{MEGA})$$

$$BR(\tau \rightarrow \mu + \gamma) < 6.8 \times 10^{-8} \quad (\text{BABAR})$$

PSI: $BR(\mu \rightarrow e + \gamma) < 10^{-13} - 10^{-14} \quad (\text{MEG})$

Plans: $BR(\tau \rightarrow \mu + \gamma) < 10^{-8} - 10^{-9} \quad (\text{LHC})$

Standard Theory with $m_\nu \neq 0$, $U_{\text{PMNS}} \neq \mathbf{1}$:

$$BR(\mu \rightarrow e + \gamma) \sim 2.5 \times 10^{-4} \left(\frac{m_j}{M_W}\right)^4 \lesssim 10^{-46}$$

S.T.P, 1977;

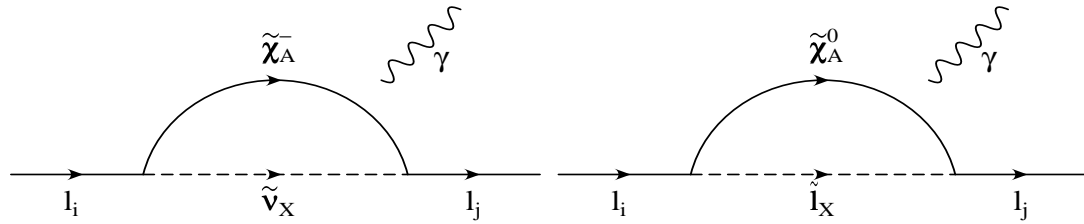
SUSY GUTs with see-saw mechanism:

- Flavour universality of SUSY breaking at $M_X \sim 10^{16}$ GeV (scalar masses m_0 , trilinear couplings $A_0 \equiv a_0 m_0$, gaugino masses $m_{1/2}$)

- RG running: “large” LFV corrections to the slepton masses $\delta m^2 \sim Y_\nu^\dagger Y_\nu$

F. Borzumati, A. Masiero, 1986.

$$(m_{sL}^2)_{ji} \approx -\frac{1}{8\pi^2} (3m_0^2 + A_0^2) (Y_\nu^\dagger Y_\nu)_{ji} \log \frac{M_X}{M_R}, \quad m_0, A_0 - \text{ at } M_X .$$



$$BR(l_i \rightarrow l_j + \gamma) \simeq \alpha^3 \left(\frac{(3 + A_0^2) m_0^2}{8\pi^2 m_S^4 G_F} \right)^2 |(Y_\nu^\dagger L Y_\nu)_{ij}|^2 \tan^2 \beta \sim \frac{|(Y_\nu^\dagger L Y_\nu)_{ij}|^2}{G_F^2 m_{\text{SUSY}}^4} \tan^2 \beta,$$

m_S - sparticle mass (at M_X); $L_k = \ln M_X/M_k$, $k = 1, 2, 3$.

$$m_S^8 \simeq 0.5 m_0^2 m_{1/2}^2 (m_0^2 + 0.6 m_{1/2}^2)^2.$$

S.T.P., S. Profumo, Y. Takahashi, C. Yaguna, 2003

For, e.g., $A_0 \sim m_S \sim m_0 \sim 10^2$ GeV, $M_X \sim 10^{16}$ GeV, $M_R \sim 10^{11}$ GeV, $\tan \beta = 10$,

$$BR(l_i \rightarrow l_j + \gamma) \sim 10^{-11} \left| (m_D^\dagger m_D)_{ij} \text{ GeV}^{-2} \right|^2$$

“Benchmark SUSY scenario”:

$$m_0 = m_{1/2} = 250 \text{ GeV}, \quad A_0 = a_0 m_0 = -100 \text{ GeV}, \quad \tan \beta \sim 5 - 10.$$

$$\chi_1^0: m_1 \sim 100 \text{ GeV}; \quad \chi_2^0, \chi^+: m_2 \sim 250 \text{ GeV}; \quad m_{sq} \sim (400 - 600) \text{ GeV (LHC)}.$$

$$BR(l_i \rightarrow l_j + \gamma) \simeq 9.1 \times 10^{-10} \left| (Y_\nu^\dagger L Y_\nu)_{ij} \right|^2 \tan^2 \beta, \quad \tan^2 \beta \gtrsim 10.$$

- Leptogenesis: $\text{Im}(Y_\nu Y_\nu^\dagger)^2$, $(Y_\nu Y_\nu^\dagger)_{11}$
- $\mu \rightarrow e + \gamma$, etc.: $Y_\nu^\dagger L Y_\nu$
- See-saw: $Y_\nu = \sqrt{D_N} R \sqrt{D_\nu} (U_{\text{PMNS}})^\dagger$, all at M_R ; $R^T R = \mathbf{1}$.

J.A. Casas and A. Ibarra, 2001

Leptogenesis: $Y_\nu Y_\nu^\dagger = \sqrt{D_N} R D_\nu R^\dagger \sqrt{D_N}$, R should be complex.

$R = R_{12}(\omega_{12})R_{13}(\omega_{13})R_{23}(\omega_{23}) = R_{12}(\omega_{12})R_{23}(\omega_{23})R_{12}(\omega'_{12})$, ω_{ij} -complex.

$\mu \rightarrow e + \gamma$: $(Y_\nu^\dagger L Y_\nu)_{21} = (Y_\nu^\dagger)_{21} L_1 (Y_\nu)_{11} + (Y_\nu^\dagger)_{22} L_2 (Y_\nu)_{21} + (Y_\nu^\dagger)_{23} L_3 (Y_\nu)_{31}$

$(Y_\nu^\dagger)_{21} L_1 (Y_\nu)_{11} \propto \sqrt{m_j m_k} M_1 / v_u^2$, $j, k \neq 1$ (3), NH (IH);

$(Y_\nu^\dagger)_{22} L_2 (Y_\nu)_{21} \propto \sqrt{m_j m_k} M_2 / v_u^2$;

$(Y_\nu^\dagger)_{23} L_3 (Y_\nu)_{31} \propto \sqrt{m_j m_k} M_3 / v_u^2$.

Hierarchical spectrum: $M_1 \ll M_2 \ll M_3$

Low Energy Leptonic CPV and Leptogenesis

Assume:

$$10^9 \lesssim M_1 \lesssim 10^{12} \text{ GeV}, \quad M_1 \ll M_2, M_3$$

The “one-flavor” approximation used is correct only if $Y_{e,\mu,\tau}$ - “small”:

Boltzmann eqn. for $n(N_1)$ and $\Delta L = \Delta(L_e + L_\mu + L_\tau)$.

$Y_l \bar{L}_l \Phi l_R$ - out of equilibrium at $T \sim M_1$.

At $M_1 \sim 10^{12}$ GeV: Y_τ - in equilibrium, $Y_{e,\mu}$ - not.

At $M_1 \sim 10^9$ GeV: Y_τ, Y_μ - in equilibrium, Y_e - not.

Thus, at $M_1 \sim 10^{12}$ GeV: $L_\tau, \Delta L_\tau$ - distinguishable;

$L_e, L_\mu, \Delta L_e, \Delta L_\mu$ - individually not distinguishable;

$L_e + L_\mu, \Delta(L_e + L_\mu)$

A. Abada et al., 2006; E. Nardi et al., 2006

A. Abada et al., 2006;

S. Blanchet and P. Di Bari, 2006.

Individual asymmetries:

$$\varepsilon_{1\alpha} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left(\sum_{\beta\rho} m_\beta^{1/2} m_\rho^{3/2} U_{\alpha\beta}^* U_{\alpha\rho} R_{1\beta} R_{1\rho}^* \right)}{\sum_{\beta} m_\beta |R_{1\beta}|^2}$$

$$Y_\alpha \sim \varepsilon_{1\alpha} \kappa(\widetilde{m}_{1\alpha}), \quad \widetilde{m}_{1\alpha} = |Y_{\nu 1\alpha}|^2 v^2 / M_1$$

The baryon asymmetry is

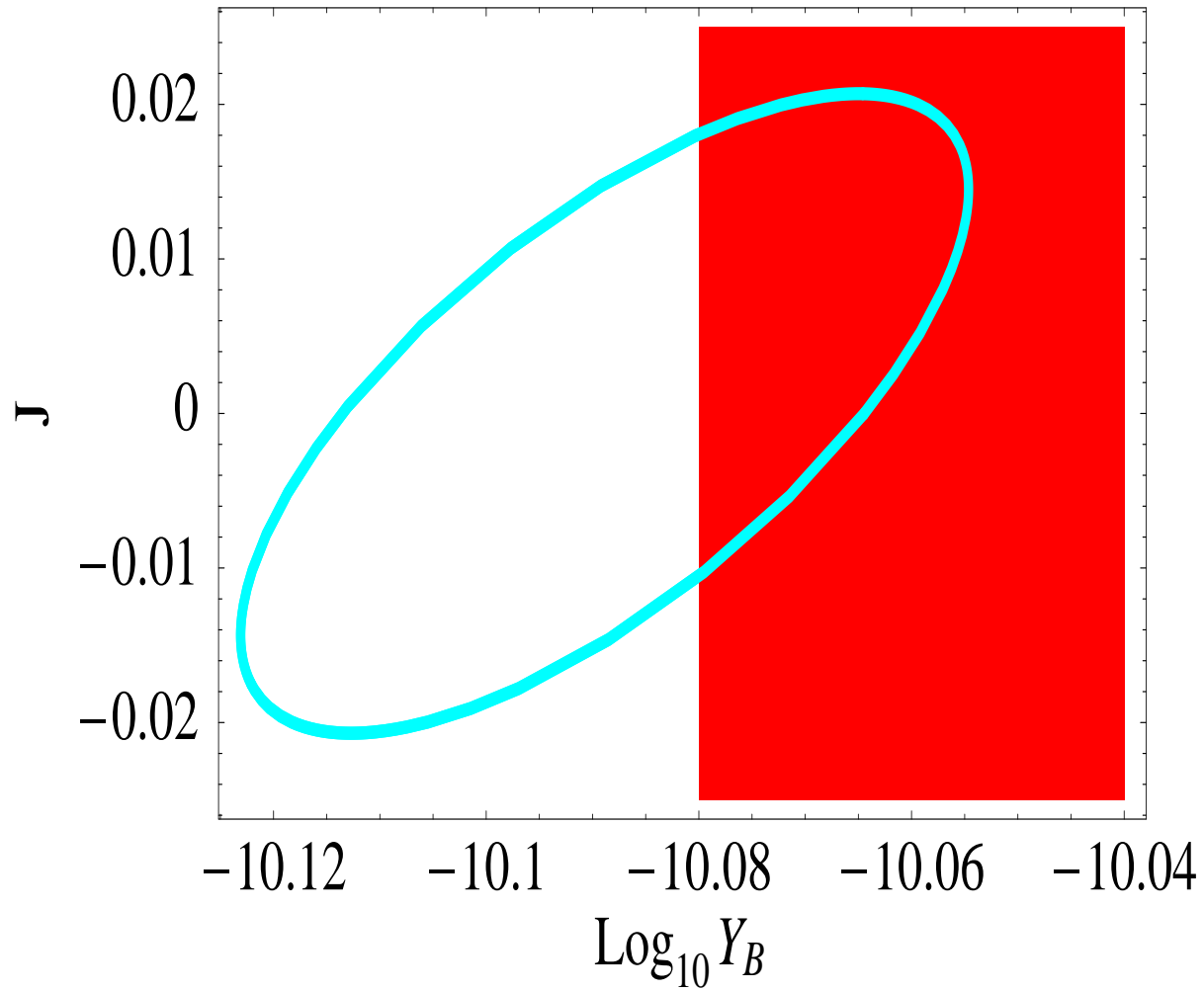
$$Y_B = -(12/37) \left((115/36) Y_2 + (37/9) Y_\tau \right),$$

$$Y_2 = Y_{e+\mu}, \quad \varepsilon_2 = \varepsilon_{1e} + \varepsilon_{1\mu}, \quad \widetilde{m}_2 = \widetilde{m}_{1e} + \widetilde{m}_{1\mu}$$

Real R : $\varepsilon_{1\alpha} \neq 0$, CPV from U

Example: NH spectrum, $M_1 \ll M_2 \ll M_3$,

$$\varepsilon_\tau \simeq \frac{3M_1}{16\pi v^2} \frac{(\Delta m_\odot^2 \Delta m_{31}^2)^{1/4} R_{12} R_{13}}{\sqrt{\Delta m_\odot^2 / \Delta m_{31}^2 R_{12}^2 + R_{13}^2}} \times \left(c_{12} c_{23} s_{23} \sin \frac{\alpha_{32}}{2} - s_{12} c_{23}^2 s_{13} \sin \left(\delta - \frac{\alpha_{32}}{2} \right) \right)$$



$M_1 = 10^{11}$ GeV, $R_{12} = 0.95$, $R_{13} = 0.3$, $\alpha_{32} = \pi$, $s_{13} = 0.09$, $s_{23} = 1/\sqrt{2}$,
 $s_{12} = 1/\sqrt{3}$, $0 \leq \delta \leq 2\pi$; $J \equiv \text{Im}(U_{e1}U_{e2}^*U_{\mu1}^*U_{\mu2})$.

S. Pascoli, S.T.P., A. Riotto, 2006.

The Role of LFV Decays: $\mu \rightarrow e + \gamma$

Assume:

$M_1 \ll M_2 \ll M_3$, $M_3 \gtrsim 5 \times 10^{13}$ GeV (GUTs)
($m_\nu \cong m_D^2/M_R$; $m_D \sim 175$ GeV, $m_\nu \sim 5 \times 10^{-2}$ eV, then $M_R \sim 6 \times 10^{14}$ GeV.)

$M_{SUSY} \sim (100 - 600)$ GeV (LHC), e.g.,

$$m_0 = m_{1/2} = 250 \text{ GeV}, \quad A_0 = a_0 m_0 = -100 \text{ GeV},$$

$BR(\mu \rightarrow e + \gamma) < 1.2 \times 10^{-11}$ implies:

terms $\sim M_3$ in $|(Y_\nu^\dagger L Y_\nu)_{21}|$ - **suppressed**, i.e., $Y_{\nu 32} \cong 0$, or $Y_{\nu 31} \cong 0$.

One possible solution - the form of **R**:

$$R = R_{12}(\omega_{12})R_{13}(\pi/2)R_{23}(\omega_{23}) = R(\omega_{12} - \omega_{23}) \equiv R(\omega), \quad \text{NH},$$

$$R = R_{12}(\omega_{12})R_{23}(0)R_{12}(\omega'_{12}) = R_{12}(\omega_{12} + \omega'_{12}) \equiv R_{12}(\omega), \quad \text{IH}.$$

NH spectrum:

$$\mathbf{R} \simeq \begin{pmatrix} 0 & \sin \omega & \cos \omega \\ 0 & \cos \omega & -\sin \omega \\ -1 & 0 & 0 \end{pmatrix}, \quad \omega = \rho + i\sigma. \quad (6)$$

S.T.P., W. Rodejohann, T. Shindou, Y. Takanishi, 2005

also: J. Ellis et al., 2004; A. Ibarra, G.G. Ross, 2004

The terms $\sim M_2$ in $|(Y_\nu^\dagger L Y_\nu)_{21}|$ – **dominant**.

Leptogenesis, NH spectrum: $M_1 \gtrsim 10^{10}$ GeV, ω –complex.

$M_2 \gtrsim \times 10^{11}$ GeV: predicted $BR(\mu \rightarrow e + \gamma) \sim 10^{-12}$

$$\epsilon_1 \simeq -\frac{3}{8\pi} \left(\frac{m_3 M_1}{v_u^2} \right) \frac{\text{Im} \left[c_\omega^2 + \frac{\Delta m_\odot^2}{\Delta m_{31}^2} s_\omega^2 \right]}{|c_\omega|^2 + \sqrt{\frac{\Delta m_\odot^2}{\Delta m_{31}^2}} |s_\omega|^2}$$

$$|\epsilon_1| \lesssim \frac{3}{8\pi} \left(\frac{m_3 M_1}{v_u^2} \right) \simeq 1.97 \times 10^{-7} \left(\frac{m_3}{0.05 \text{ eV}} \right) \left(\frac{M_1}{10^9 \text{ GeV}} \right) \left(\frac{174 \text{ GeV}}{v_u} \right)^2.$$

$$\tilde{m}_1 \simeq m_3 |c_\omega|^2 + m_2 |s_\omega|^2 = \frac{1}{2}(m_3 + m_2) \cosh 2\sigma + \frac{1}{2}(m_3 - m_2) \cos 2\rho \geq m_2.$$

$$9 \times 10^{-3} \text{ eV} < \tilde{m}_1 \lesssim 0.12 \text{ eV} : \quad 1.9 \times 10^{-3} \lesssim \kappa < 3.9 \times 10^{-2}$$

IH (QD) spectrum:

$$\mathbf{R} \simeq \begin{pmatrix} \cos \omega & \sin \omega & 0 \\ -\sin \omega & \cos \omega & 0 \\ 0 & 0 & 1 \end{pmatrix}. \quad (7)$$

The terms $\sim M_2$ in $|(\mathbf{Y}_\nu^\dagger L \mathbf{Y}_\nu)_{21}|$ – **dominant**.

Leptogenesis, IH spectrum: $M_1 \gtrsim 7 \times 10^{12}$ GeV, ω –complex.

$M_2 \gtrsim 5 \times 10^{13}$ GeV: predicted $BR(\mu \rightarrow e + \gamma) \gg 10^{-11}$

$$\epsilon_1 \simeq -\frac{3}{8\pi} \left(\frac{m_2 M_1}{v_u^2} \right) \frac{\Delta m_\odot^2}{|\Delta m_{31}^2|} \frac{\text{Im} [\sin^2 \omega_{12}]}{\left(1 + \frac{\Delta m_\odot^2}{2|\Delta m_{31}^2|} \right) |\sin \omega_{12}|^2 + |\cos \omega_{12}|^2}$$

$$|\epsilon_1| \lesssim \frac{3}{16\pi} \left(\frac{m_2 M_1}{v_u^2} \right) \frac{\Delta m_\odot^2}{|\Delta m_{31}^2|} \simeq 3.2 \times 10^{-9} \left(\frac{m_2}{0.05 \text{ eV}} \right) \left(\frac{M_1}{10^9 \text{ GeV}} \right) \left(\frac{174 \text{ GeV}}{v_u} \right)^2,$$

$$\tilde{m}_1 \simeq m_{1,2} (|\cos \omega_{12}|^2 + |\sin \omega_{12}|^2) = m_{1,2} \cosh 2\sigma \geq m_{1,2}.$$

$$5 \times 10^{-2} \text{ eV} < \tilde{m}_1 \lesssim 0.1 \text{ eV} : 2.4 \times 10^{-3} \lesssim \kappa < 5.4 \times 10^{-3}$$

Two possibilities:

- $M_{SUSY} \sim (600 - 2000) \text{ GeV}$,
($m_{1/2} \gg m_0$, e.g, $m_0 = 300 \text{ GeV}$, $m_{1/2} = 1400 \text{ GeV}$, $a_0 m_0 = 0$)
- $M_{SUSY} \sim (100 - 600) \text{ GeV}$, but $\mathbf{Y}_{\nu 21} = 0$, or $\mathbf{Y}_{\nu 22} = 0$

A. $\mathbf{Y}_{\nu 21} = 0$:

$$\tan \omega = e^{-i\alpha/2} \tan \theta_{12}.$$

B. $\mathbf{Y}_{\nu 22} \cong 0$, neglecting s_{13} :

$$\tan \omega = -e^{-i\alpha/2} \cot \theta_{12}.$$

Leptogenesis: ω -complex; thus $\alpha \neq 0, \pi$, CP-violating values

B. $\mathbf{Y}_{\nu 22} = 0$, including s_{13} :

$$\tan \omega = -\frac{c_{12} - s_{12}s_{13}e^{-i\delta}}{s_{12} + c_{12}s_{13}e^{-i\delta}} e^{-i\alpha/2}$$

IH spectrum:

$$|\langle m \rangle| \cong \sqrt{\Delta m_{13}^2} |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|$$

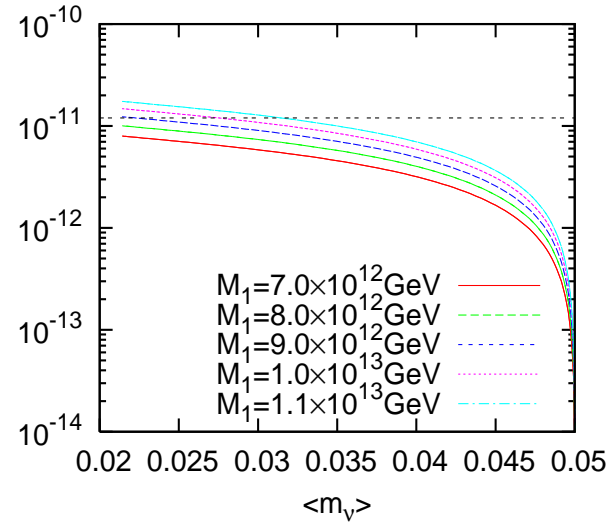
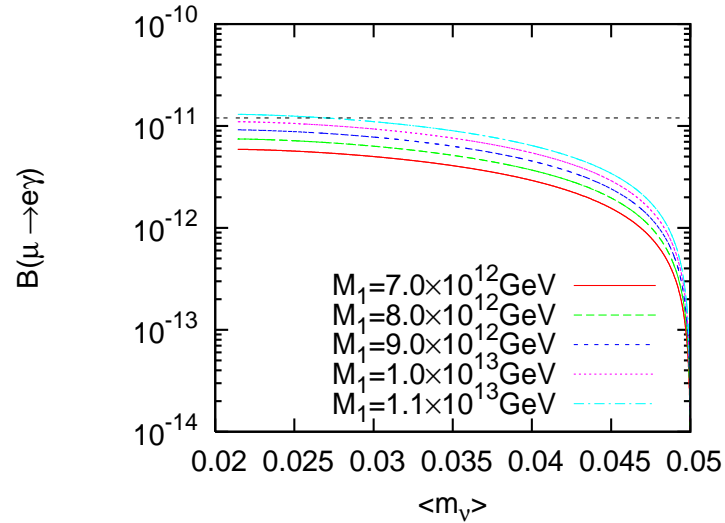
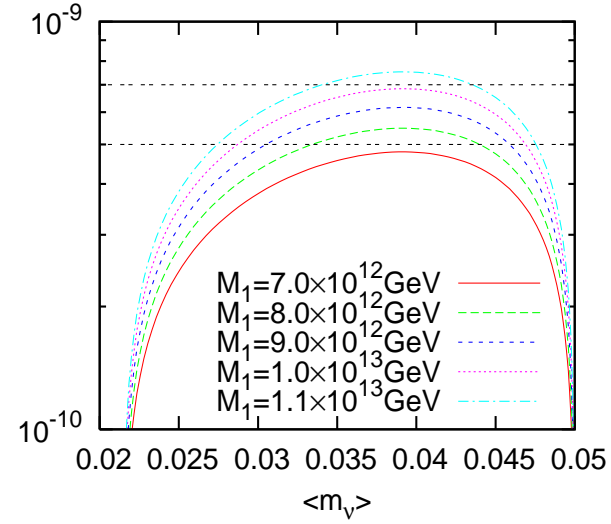
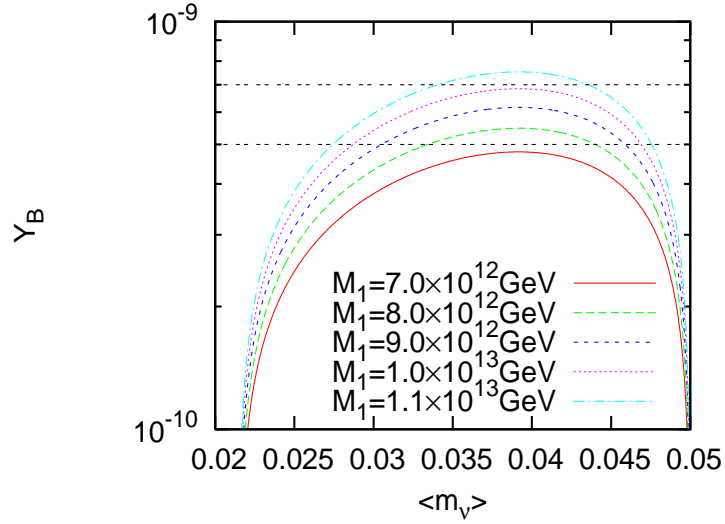


Figure 6: Predicted values of Y_B and $B(\mu \rightarrow e\gamma)$ for $s_{13} = 0$. The SUSY parameters are fixed as $m_0 = m_{1/2} = 450$, $A_0 = 0$, and $\tan\beta = 5$.

Conclusions

Determining the nature - Dirac or Majorana, of massive neutrinos is of fundamental importance for understanding the origin of neutrino masses.

The see-saw mechanism provides a link between ν -mass generation and BAU.

SUSY see-saw: LFV processes $\mu \rightarrow e + \gamma$, etc.

LHC: constraints on (discovery of?) SUSY.

$\mu \rightarrow e + \gamma$, leptogenesis - significant constraints on the theory.

Majorana CPV phases in U_{PMNS} : $(\beta\beta)_{0\nu}$ -decay, Y_{B} , $\mu \rightarrow e + \gamma$.

CPV phases in U_{PMNS} can be the leptogenesis CPV parameters.

Low energy leptonic CPV can be directly related to the existence of BAU.

The searches for leptonic CP-violation are of fundamental importance.

Supporting Slide

Oscillation Parameters

$$\Delta m_{\odot}^2 = 8.0 \times 10^{-5} \text{ eV}^2, \quad 3\sigma(\Delta m_{\odot}^2) = 12\%,$$

$$\sin^2 \theta_{\odot} = 0.30, \quad 3\sigma(\sin^2 \theta_{\odot}) = 24\%,$$

$$|\Delta m_{\text{atm}}^2| = 2.5 \times 10^{-3} \text{ eV}^2, \quad 3\sigma(|\Delta m_{\text{atm}}^2|) = 28\%.$$

Future:

3 kTy KamLAND: $3\sigma(\Delta m_{\odot}^2) = 7\%$, $3\sigma(\sin^2 \theta_{\odot}) = 18\%$;

A. Bandyopadhyay et al., hep-ph/0410283

SK-Gd (0.1% Gd: 43×(KL $\bar{\nu}_e$ rate)), 3y: $3\sigma(\Delta m_{\odot}^2) \cong 4\%$

S. Choubey, S.T.P., hep-ph/0404103;

J. Beacom and M. Vagins, hep-ph/0309300

KL type reactor $\bar{\nu}_e$ detector, $L \sim 60$ km, ~ 60 GW kTy: $3\sigma(\sin^2 \theta_{\odot}) \cong 12\%$

A. Bandyopadhyay et al., hep-ph/0410283 and hep-ph/0302243;

H. Minakata et al., hep-ph/0407326

T2K (SK): $3\sigma(|\Delta m_{\text{atm}}^2|) \cong 6\%$

$\text{sgn}(\Delta m_{\text{atm}}^2)$: ν_{atm} experiments, studying the subdominant $\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}$ and $\bar{\nu}_{\mu(e)} \rightarrow \bar{\nu}_{e(\mu)}$ oscillations; LBL ν -oscillation experiments (T2K, NO ν A); ν -factory.

$\sin^2 \theta_{13}$: reactor $\bar{\nu}_e$ experiments, $L \sim (1 - 2)$ km: Double CHOOZ, Daya-Bay, KASKA - factor (5 - 10).

Majorana Neutrinos

Can be defined in QFT using fields or states.

Fields: $\chi_k(x)$ - 4 component (spin 1/2), complex, m_k

Majorana condition:

$$C (\bar{\chi}_k(x))^T = \xi_k \chi_k(x), \quad |\xi_k|^2 = 1$$

- Invariant under proper Lorentz transformations.
- Reduces by 2 the number of components in $\chi_k(x)$.

Implications:

$$U(1) : \chi_k(x) \rightarrow e^{i\alpha} \chi_k(x) - \text{impossible}$$

- $\chi_k(x)$ cannot absorb phases.
- $Q_{U(1)} = 0 : Q_{el} = 0, L_l = 0, L = 0, \dots$
- $\chi_k(x)$: 2 spin states of a spin 1/2 absolutely neutral particle
- $\chi_k \equiv \bar{\chi}_k$

Propagators: $\Psi(x)$ –Dirac, $\chi(x)$ –Majorana

$$\langle 0|T(\Psi_\alpha(x)\bar{\Psi}_\beta(y))|0 \rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\Psi_\alpha(x)\Psi_\beta(y))|0 \rangle = 0 , \quad \langle 0|T(\bar{\Psi}_\alpha(x)\bar{\Psi}_\beta(y))|0 \rangle = 0 .$$

$$\langle 0|T(\chi_\alpha(x)\bar{\chi}_\beta(y))|0 \rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\chi_\alpha(x)\chi_\beta(y))|0 \rangle = -\xi^* S_{\alpha\kappa}^F(x-y)C_{\kappa\beta} ,$$

$$\langle 0|T(\bar{\chi}_\alpha(x)\bar{\chi}_\beta(y))|0 \rangle = \xi C_{\alpha\kappa}^{-1}S_{\kappa\beta}^F(x-y)$$

$$U_{CP} \chi(x) U_{CP}^{-1} = \eta_{CP} \gamma_0 \chi(x') , \quad \eta_{CP} = \pm i .$$

Currents of Majorana Fields